

Problem Set 1
Physics 498TBP / Spring 2002
Solutions

Problem 1: Optical Properties of Ring of BChls

(a) Let us first construct the Hamiltonian in the basis of $\{|\alpha\rangle : \alpha = 1, \dots, 2N\}$ where $N = 8$. The diagonal elements are assumed to be the same. The nearest-neighbor interaction energies are also the same because of the rotational symmetry of the ring structure. Thus, the Hamiltonian has the form of

$$H = \begin{pmatrix} \epsilon_0 & v & & & & & & v \\ v & \epsilon_0 & v & & & & & \\ & v & \epsilon_0 & & & & & \\ & & & \ddots & & & & \\ & & & & \ddots & & & \\ & & & & & \epsilon_0 & v & \\ v & & & & & v & \epsilon_0 & \end{pmatrix}. \quad (1)$$

The diagonal elements are $\epsilon_0 = 1.6 \text{ eV}$. We now calculate v :

$$v = \langle 1|H|2\rangle = \frac{\vec{d}_1 \cdot \vec{d}_2}{r_{12}^3} - \frac{3(\vec{r}_{12} \cdot \vec{d}_1)(\vec{r}_{12} \cdot \vec{d}_2)}{r_{12}^5}. \quad (2)$$

Here,

$$r_{12} = 2 \cdot 25 \text{ \AA} \cdot \sin(\pi/2N) = 9.75 \text{ \AA} \quad (3)$$

$$\vec{d}_1 \cdot \vec{d}_2 = -d_0^2 \cos(\pi/N) = -92.4 \text{ Debye}^2 \quad (4)$$

$$(\vec{r}_{12} \cdot \vec{d}_1)(\vec{r}_{12} \cdot \vec{d}_2) = -r_{12}^2 d_0^2 \cos^2(\pi/2N) = -9152.9 \text{ Debye}^2 \text{ \AA}^2. \quad (5)$$

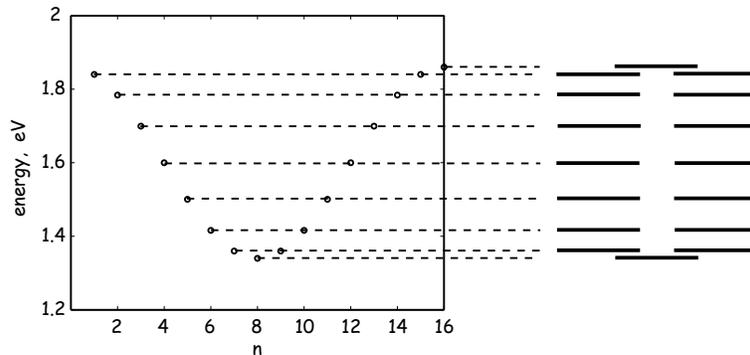
Plugging these values into Eq. 2 we find $v = 0.13 \text{ eV}$.

In class we have learned that the eigenstates (stationary states) of the Hamiltonian (Eq. 1) can be written as

$$|\tilde{n}\rangle = \frac{1}{\sqrt{2N}} \sum_{\alpha=1}^{2N} e^{in\alpha\pi/N} |\alpha\rangle, \quad n = 1, \dots, 2N. \quad (6)$$

The corresponding energies are

$$\epsilon_n = \epsilon_0 + 2v \cos \frac{n\pi}{N} = (1.6 + 2 \cdot 0.13 \cdot \cos \frac{n\pi}{8}) \text{ eV}. \quad (7)$$



(b) Transition dipole moments:

$$\begin{aligned}
\langle 0|\vec{r}|\tilde{n}\rangle &= \frac{1}{4} \sum_{\alpha=1}^{16} e^{in\alpha\pi/8} \langle 0|\vec{r}|\alpha\rangle \\
&= \frac{1}{4} \sum_{\alpha=1}^{16} e^{in\alpha\pi/8} \vec{d}_{\alpha}/e \\
&= \frac{1}{4} \sum_{\alpha=1}^{16} e^{in\alpha\pi/8} \frac{d_0}{e} \left(\hat{x} \cos \frac{9\pi\alpha}{8} + \hat{y} \sin \frac{9\pi\alpha}{8} \right), \tag{8}
\end{aligned}$$

where \hat{x} and \hat{y} are the unit vectors in x and y direction, respectively. Evaluating this formula for $n = 1, \dots, 16$, we find

$$\langle 0|\vec{r}|\tilde{n}\rangle = \begin{cases} (d_0/e)(2\hat{x} - 2i\hat{y}) = (20/e)(\hat{x} - i\hat{y}) \text{ Debye}, & n = 7 \\ (d_0/e)(2\hat{x} + 2i\hat{y}) = (20/e)(\hat{x} + i\hat{y}) \text{ Debye}, & n = 9 \\ 0, & \text{otherwise.} \end{cases} \tag{9}$$

Transition rates can be calculated from the transition dipole moments through the formula

$$k_{0 \rightarrow \tilde{n}} = N_{\omega_n} \frac{4e^2\omega_n^3}{3c^3\hbar} |\langle 0|\vec{r}|\tilde{n}\rangle|^2, \tag{10}$$

where N_{ω} is the number of photons of energy $\hbar\omega$. Using the energy spectrum obtained in Eq. 7, we find

$$k_{0 \rightarrow \tilde{n}} = \begin{cases} 0.32 N_{\omega_7}/\text{ns}, & n = 7 \\ 0.32 N_{\omega_9}/\text{ns}, & n = 9 \\ 0, & \text{otherwise.} \end{cases} \tag{11}$$

In fact, $k_{0 \rightarrow \tilde{7}} = k_{0 \rightarrow \tilde{9}}$ because $\omega_7 = \omega_9$.

Transition rate for individual BChls can be calculated in a similar way:

$$k_{0 \rightarrow \alpha} = N_{\omega_0} \frac{4e^2\omega_0^3}{3c^3\hbar} |\langle 0|\vec{r}|\alpha\rangle|^2. \tag{12}$$

Here $\hbar\omega_0 = \epsilon_0 = 1.6 \text{ eV}$ and $|\langle 0|\vec{r}|\alpha\rangle|^2 = |\vec{d}_{\alpha}/e|^2 = (100/e^2) \text{ Debye}$. Therefore,

$$k_{0 \rightarrow \alpha} = 0.065 N_{\omega_0}/\text{ns}. \tag{13}$$

Assuming $N_{\omega_{7,9}} \approx N_{\omega_0}$,

$$\frac{k_{0 \rightarrow \tilde{7}, \tilde{9}}}{k_{0 \rightarrow \alpha}} \approx 4.9. \tag{14}$$

Problem 2: Semiclassical Theory of Electron Transfer

(a) $p_0(q)$ is the Boltzmann distribution corresponding to $V_r(q) = fq^2/2$:

$$p_0(q) = \sqrt{\frac{\beta f}{2\pi}} e^{-\beta fq^2/2}, \tag{15}$$

where the prefactor was determined by the normalization condition, namely $\int_{-\infty}^{\infty} dq p_0(q) = 1$. By inverting

$$E(q) = V_p(q) - V_r(q) = f(q - q_0)^2/2 + E_0 - fq^2/2, \tag{16}$$

we find

$$q(E) = \frac{1}{fq_0} \left(\frac{1}{2}fq_0^2 + E_0 - E \right) \quad (17)$$

and

$$\left| \frac{dq}{dE} \right| = \frac{1}{fq_0}. \quad (18)$$

Therefore,

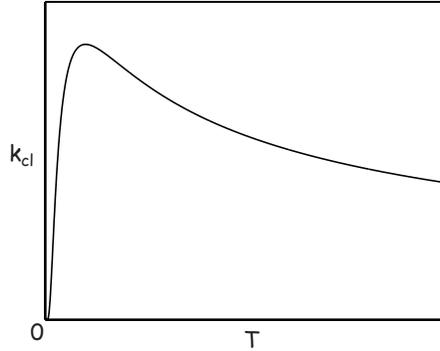
$$S_{\text{cl}}(E) = p_0[q(E)] \left| \frac{dq}{dE} \right| \quad (19)$$

$$= \sqrt{\frac{\beta f}{2\pi}} \exp \left[-\frac{\beta f}{2(fq_0)^2} \left(E_0 - E + \frac{1}{2}fq_0^2 \right)^2 \right] \frac{1}{fq_0} \quad (20)$$

$$= \frac{1}{\sqrt{2\pi\sigma_{\text{cl}}f^2q_0^2}} \exp \left[-\frac{(E_0 - E + fq_0^2/2)^2}{2f^2q_0^2\sigma_{\text{cl}}} \right]. \quad (21)$$

(b)

$$k_{\text{cl}} = \frac{2\pi}{\hbar} |U|^2 S_{\text{cl}}(0) = \frac{2\pi}{\hbar} |U|^2 \frac{1}{\sqrt{2\pi\sigma_{\text{cl}}f^2q_0^2}} \exp \left[-\frac{(E_0 + fq_0^2/2)^2}{2f^2q_0^2\sigma_{\text{cl}}} \right]. \quad (22)$$



(c) The density operator is

$$\rho_0 = \frac{1}{Z} e^{-\beta H_r}, \quad (23)$$

where Z is the partition function. In the basis of energy eigenstates,

$$[\rho_0]_{nm} = \langle \tilde{n} | \frac{1}{Z} e^{-\beta H_r} | \tilde{m} \rangle = \frac{1}{Z} \langle \tilde{n} | \tilde{m} \rangle e^{-\beta \hbar \omega (m+1/2)} = \frac{1}{Z} \delta_{nm} e^{-\beta \hbar \omega (n+1/2)}. \quad (24)$$

The partition function Z can be determined by the normalization condition:

$$1 = \text{tr} \rho_0 = \sum_{n=0}^{\infty} [\rho_0]_{nn} = \frac{1}{Z} \sum_{n=0}^{\infty} e^{-\beta \hbar \omega n} e^{-\beta \hbar \omega / 2} = \frac{1}{Z} e^{-\beta \hbar \omega / 2} (1 - e^{-\beta \hbar \omega})^{-1}. \quad (25)$$

Therefore,

$$Z = (e^{\beta \hbar \omega / 2} - e^{-\beta \hbar \omega / 2})^{-1}, \quad (26)$$

and

$$[\rho_0]_{nm} = \delta_{nm} (1 - e^{-\beta \hbar \omega}) e^{\beta \hbar \omega / 2} e^{-\beta \hbar \omega (n+1/2)}. \quad (27)$$

(d)

$$p_{\text{qm}}(q') = \text{tr}[\rho_0 \delta(q - q')] = \text{tr} \left[\rho_0 \frac{1}{2\pi} \int_{-\infty}^{\infty} dx e^{ix(q-q')} \right] = \frac{1}{2\pi} \int_{-\infty}^{\infty} dx e^{-ixq'} \Phi(x) \quad (28)$$

$$\Phi(x) \equiv \text{tr}[\rho_0 e^{ixq}] = \langle e^{ixq} \rangle_0, \quad (29)$$

where $\langle \cdot \rangle_0$ denotes the ensemble average with respect to the density operator ρ_0 . We now turn to the second-quantization representation:

$$q = \sqrt{\frac{\hbar}{2m\omega}} (a + a^\dagger), \quad (30)$$

where a is the lowering operator and a^\dagger is the raising operator. Introducing $\xi \equiv (\hbar/2m\omega)^{1/2}$ for simplicity, we have

$$\Phi(x) = \langle e^{ix\xi(a+a^\dagger)} \rangle_0. \quad (31)$$

Using the identity,

$$e^A e^B = e^{A+B+[A,B]/2} \quad \text{if } [A, B] \text{ is a complex number}, \quad (32)$$

we factorize the exponential function:

$$\Phi(x) = e^{-x^2\xi^2/2} \langle e^{ix\xi a^\dagger} e^{ix\xi a} \rangle_0 = e^{-x^2\xi^2/2} \left\langle \sum_{m=0}^{\infty} \frac{1}{m!} (ix\xi a^\dagger)^m \sum_{n=0}^{\infty} \frac{1}{n!} (ix\xi a)^n \right\rangle_0, \quad (33)$$

where exponential functions were Taylor-expanded in the last step. Among the possible pairs of (m, n) , only those satisfying $m = n$ survive the ensemble average:

$$\Phi(x) = e^{-x^2\xi^2/2} \sum_{n=0}^{\infty} \frac{1}{n!n!} (ix\xi)^{2n} \langle a^{\dagger n} a^n \rangle_0. \quad (34)$$

By Wick's theorem, we have

$$\langle a^{\dagger n} a^n \rangle_0 = n! \langle a^\dagger a \rangle_0^n, \quad (35)$$

where the ensemble average $\langle a^\dagger a \rangle_0$ can be calculated in the energy eigenbasis:

$$\begin{aligned} \langle a^\dagger a \rangle_0 &= \sum_{m=0}^{\infty} \langle \tilde{m} | \frac{1}{Z} e^{-\beta H_r} a^\dagger a | \tilde{m} \rangle \\ &= \frac{1}{Z} \sum_m e^{-\beta \hbar \omega (m+1/2)} m = \frac{1}{Z} \sum_m e^{-\beta \hbar \omega (m+1/2)} \left(m + \frac{1}{2} - \frac{1}{2} \right) \\ &= \frac{1}{Z} \frac{\partial}{\partial (-\beta \hbar \omega)} \sum_m e^{-\beta \hbar \omega (m+1/2)} - \frac{1}{2Z} \sum_m e^{-\beta \hbar \omega (m+1/2)} \\ &= \frac{1}{Z} \frac{\partial}{\partial (-\beta \hbar \omega)} Z - \frac{1}{2} = \frac{1}{2} \left(\coth \frac{\beta \hbar \omega}{2} - 1 \right). \end{aligned} \quad (36)$$

In the last step we used the partition function obtained in Eq. 26. Combining Eqs. 34, 35, and 36 we find

$$\begin{aligned} \Phi(x) &= e^{-x^2\xi^2/2} \sum_{n=0}^{\infty} \frac{1}{n!} (ix\xi)^{2n} \frac{1}{2^n} \left(\coth \frac{\beta \hbar \omega}{2} - 1 \right)^n \\ &= e^{-x^2\xi^2/2} \exp \left[-\frac{x^2\xi^2}{2} \left(\coth \frac{\beta \hbar \omega}{2} - 1 \right) \right] = e^{-\sigma_{\text{qm}} x^2/2}, \end{aligned} \quad (37)$$

where

$$\sigma_{\text{qm}} = \xi^2 \coth \frac{\beta \hbar \omega}{2} = \frac{\hbar}{2m\omega} \coth \frac{\beta \hbar \omega}{2}. \quad (38)$$

(e) We calculate $p_{\text{qm}}(q')$ and then $S_{\text{qm}}(E)$. From Eq. 28,

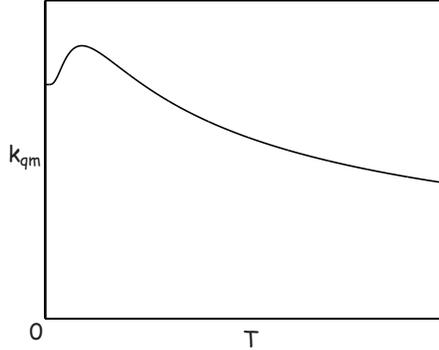
$$\begin{aligned}
p_{\text{qm}}(q') &= \frac{1}{2\pi} \int_{-\infty}^{\infty} dx e^{-ixq'} e^{-\sigma_{\text{qm}}x^2/2} \\
&= \frac{1}{2\pi} \int_{-\infty}^{\infty} dx \exp \left[-\frac{\sigma_{\text{qm}}}{2} \left(x + \frac{iq'}{\sigma_{\text{qm}}} \right)^2 - \frac{q'^2}{2\sigma_{\text{qm}}} \right] \\
&= (2\pi\sigma_{\text{qm}})^{-1/2} \exp \left[-\frac{q'^2}{2\sigma_{\text{qm}}} \right].
\end{aligned} \tag{39}$$

Combining Eqs. 17, 18, and 39 leads to

$$\begin{aligned}
S_{\text{qm}}(E) &= p_{\text{qm}}[q(E)] \left| \frac{dq}{dE} \right| \\
&= (2\pi\sigma_{\text{qm}})^{-1/2} \exp \left[-\frac{1}{2\sigma_{\text{qm}}} \frac{1}{f^2 q_0^2} (f q_0^2/2 + E_0 - E)^2 \right] \frac{1}{f q_0} \\
&= (2\pi\sigma_{\text{qm}} f^2 q_0^2)^{-1/2} \exp \left[-\frac{(E_0 - E + f q_0^2/2)^2}{2f^2 q_0^2 \sigma_{\text{qm}}} \right].
\end{aligned} \tag{40}$$

(f)

$$\begin{aligned}
k_{\text{qm}} &= \frac{2\pi}{\hbar} |U|^2 S_{\text{qm}}(0) = \frac{\sqrt{2\pi}}{\hbar} \frac{|U|^2}{\sqrt{\sigma_{\text{qm}} f^2 q_0^2}} \exp \left[-\frac{(E_0 + f q_0^2/2)^2}{2f^2 q_0^2 \sigma_{\text{qm}}} \right] \\
&= \frac{\sqrt{2\pi}}{\hbar} \frac{|U|^2}{|f q_0|} \left(\frac{\hbar}{2m\omega} \coth \frac{\hbar\omega}{2k_B T} \right)^{-1/2} \exp \left[-\frac{(E_0 + f q_0^2/2)^2}{2f^2 q_0^2} \left(\frac{\hbar}{2m\omega} \coth \frac{\hbar\omega}{2k_B T} \right)^{-1} \right].
\end{aligned} \tag{41}$$



Take $T \rightarrow 0$ limit of Eqs. 38 and 39:

$$\sigma_{\text{qm}} \rightarrow \frac{\hbar}{2m\omega} \tag{42}$$

$$p_{\text{qm}}(q') \rightarrow \left(\frac{m\omega}{\pi\hbar} \right)^{1/2} \exp \left[-\frac{m\omega q'^2}{\hbar} \right] = |\psi_0(q')|^2, \tag{43}$$

where ψ_0 is the ground-state wave function for the harmonic oscillator representing the reactant state. The quantum mechanical transition rate k_{qm} does not vanish at $T = 0$ because there exists quantum mechanical fluctuation even at zero temperature; the ground-state wave function is not a delta function.

Problem 3: End-End Reaction of One-Dimensional Polymer

(a) Let a_i be the orientation of i th segment. It takes the value of either $+1$ or -1 with probability $1/2$. The end-to-end distance x is then given as $x = b \sum_{i=1}^{2N} a_i$. Since a_i are independent random variables, we can apply the central limit theorem. The theorem states that as $N \rightarrow \infty$ the distribution for x becomes Gaussian.

In order to determine the actual formula for the Gaussian distribution, we need only the mean and the variance:

$$\langle x \rangle = b \sum_i \langle a_i \rangle = 0 \quad (44)$$

$$\langle x^2 \rangle = b^2 \sum_i \sum_j \langle a_i a_j \rangle = b^2 \sum_i \langle a_i^2 \rangle = 2Nb^2. \quad (45)$$

Here we have used the independence, namely $\langle a_i a_j \rangle = \delta_{ij} \langle a_i^2 \rangle = \delta_{ij}$. The Gaussian distribution with the above mean and variance is

$$p_0(x) = (4\pi b^2 N)^{-1/2} \exp(-x^2/4b^2 N). \quad (46)$$

(b) Any probability distribution that makes the right-hand side of the Smoluchowski equation vanish is stationary. It is straightforward to show that $p_0(x)$ makes the right-hand side vanish:

$$\begin{aligned} & (4Nb^2 \partial_x^2 + 2\partial_x x) p_0(x) \\ &= (4\pi b^2 N)^{-1/2} \left[4Nb^2 \left(-\frac{1}{2b^2 N} + \frac{x^2}{4b^4 N^2} \right) + 2 - \frac{x^2}{b^2 N} \right] \exp\left(-\frac{x^2}{4b^2 N}\right) \\ &= 0. \end{aligned} \quad (47)$$

The distribution $p_0(x)$ is therefore a stationary solution.

(c)

$$p(x, t|x_0, t_0) = [4b^2 N \pi S(t, t_0)]^{-1/2} \exp\left[-\frac{(x - x_0 e^{-2(t-t_0)/\tau})^2}{4b^2 N S(t, t_0)}\right] \quad (48)$$

$$S(t, t_0) = 1 - e^{-4(t-t_0)/\tau}. \quad (49)$$

A little algebra leads to

$$\begin{aligned} \tau \partial_t p(x, t|x_0, t_0) &= (4Nb^2 \partial_x^2 + 2\partial_x x) p(x, t|x_0, t_0) \\ &= -\frac{2b^2 N S(t, t_0) + x x_0 (e^{-2(t-t_0)/\tau} + e^{2(t-t_0)/\tau}) - (x^2 + x_0^2)}{2b^3 N^{3/2} \pi^{1/2} S(t, t_0)^{5/2} e^{4(t-t_0)/\tau}} \exp\left[-\frac{(x - x_0 e^{-2(t-t_0)/\tau})^2}{4b^2 N S(t, t_0)}\right], \end{aligned} \quad (50)$$

which means that $p(x, t|x_0, t_0)$ satisfies the Smoluchowski equation. We now show that it also satisfies the initial condition. First, it is normalized at any time t later than t_0 : $\int_{-\infty}^{\infty} dx p(x, t|x_0, t_0) = 1$. Second, as t approaches t_0 , $p(x, t|x_0, t_0)$ gets vanishingly small for $x \neq x_0$:

$$\lim_{t \rightarrow t_0} p(x, t|x_0, t_0) = \lim_{S \rightarrow 0^+} (4b^2 N \pi S)^{-1/2} \exp\left[-\frac{(x - x_0)^2}{4b^2 N S}\right] = 0. \quad (51)$$

These two properties imply that $\lim_{t \rightarrow t_0} p(x, t|x_0, t_0) = \delta(x - x_0)$. This completes the proof.

As $t \rightarrow \infty$, $S(t, t_0) \rightarrow 1$ and the solution $p(x, t|x_0, t_0)$ indeed relaxes to the equilibrium distribution $p_0(x)$.